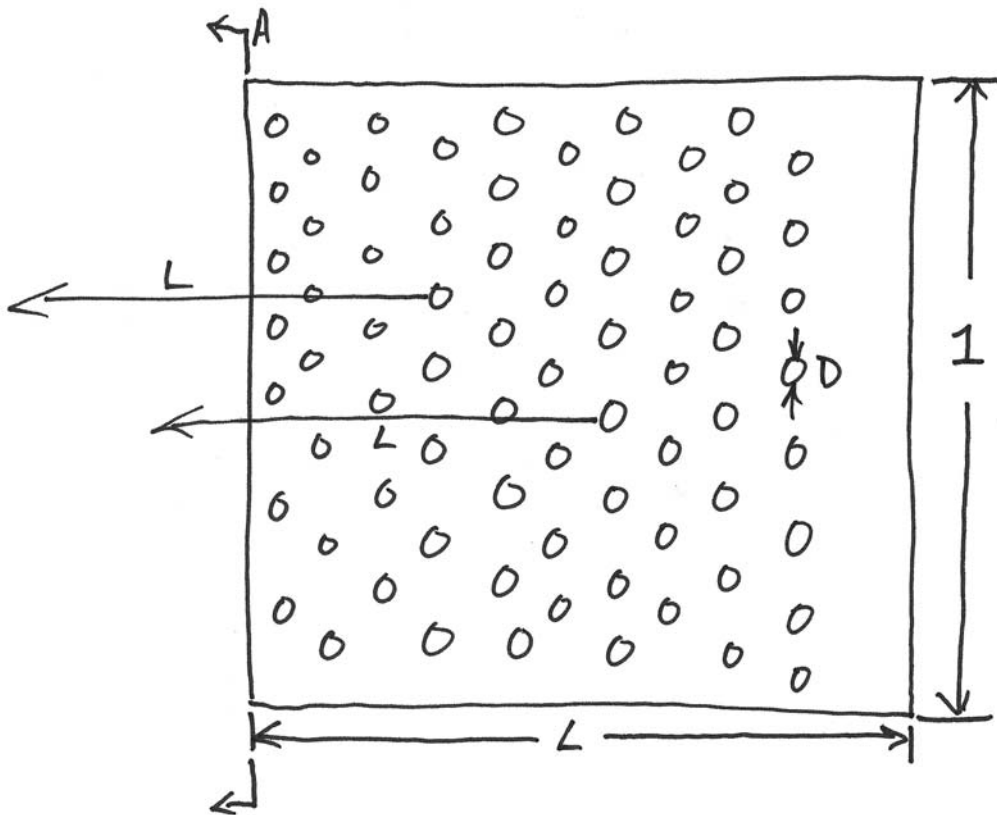


## Lecture 24—Probability Bed Load Equations

Albert Einstein claimed that he became a physicist because he found sediment transport (in which he was originally trained) too intractable a problem. How typical, then, that his son decided to one-up dad by taking the problem on.

H.A. Einstein, in 1942, tried a very different tack from the excess shear concept. Einstein described a turbulent fluid where forces vary in both space and time; the *probability* of movement of a particle depends on the *probability* that the fluid forces will exceed resistive forces.

So, let's consider a flat bed:



And let's assume that if a particle moves, it will move a distance  $L$ , which we can describe in terms of grain diameters,  $L = \lambda_0 D$ .

From before,

$$q_b = \frac{g_b}{(\rho_s - \rho)g}$$

And, from Einstein,

$$q_b = \frac{L}{A_1 D^2} \cdot \phi_s \cdot A_2 D^3 = \frac{A_2}{A_1} \lambda_0 D^2 \phi_s$$

so,

$$\frac{g_b}{(\rho_s - \rho)gD^2} \cdot \frac{A_1}{A_2} = \lambda_0 \phi_s$$

We would like  $\phi_s$  to be independent of time.

$$\phi = t \phi_s$$

Einstein assumes  $t \propto \frac{D}{w_s}$ , which is the time it takes for a grain to fall one diameter. From long ago,

$$w_s^2 = \frac{4g}{3C_D} D \left( \frac{\rho_s - \rho}{\rho} \right)$$

so

$$\frac{D}{w_s} = \frac{1}{\sqrt{\frac{4}{3}C_D}} \sqrt{\frac{D\rho}{g(\rho_s - \rho)}}$$

and

$$t = \frac{A_3}{F} \sqrt{\frac{D\rho}{g(\rho_s - \rho)}}$$

where  $F = \sqrt{\frac{4}{3}C_D}$  for spheres, and is given implicitly by Rubey (1933) to be:

$$F = \sqrt{\frac{2}{3} + \frac{36\mu^2}{gD^3\rho(\rho_s - \rho)}} - \sqrt{\frac{36\mu^2}{gD^3\rho(\rho_s - \rho)}}$$

Now we can formulate  $\wp$ .

$$\wp = \frac{A_1 A_3}{\lambda_0 A_2} \left[ \frac{1}{F} \frac{g_b}{(\rho_s - \rho)g} \sqrt{\frac{\rho}{\rho_s - \rho}} \frac{1}{g^{\frac{1}{2}} D^{\frac{3}{2}}} \right]$$

Additionally, Einstein assumes that the probability will be some function of particle weight over the average lift on the particle. So,

$$\wp = f\left(\frac{A_2 D^3 (\rho_s - \rho) g}{A_4 D^2 \rho V^2}\right)$$

Which  $V$  to choose?  $V$  is taken as the velocity at the top of the laminar boundary layer, and is expressed as  $V=11.6u^*$ . This makes our equation:

$$\phi = f\left(\frac{A_2 D^3 (\rho_s - \rho) g}{A_4 D^2 (135 g R S)}\right)$$

Simplifying,

$$\phi = f\left(\frac{A_2}{135 A_4} \cdot \frac{\rho_s - \rho}{\rho} \cdot \frac{D}{R S}\right)$$

Now, let  $A = \frac{A_1 A_3}{\lambda_0 A_2}$ ,  $B = \frac{A_2}{135 A_4}$ . This makes the whole Einstein equation:

$$A \left[ \frac{1}{F} \left( \frac{g_b}{(\rho_s - \rho) g} \sqrt{\frac{\rho}{(\rho_s - \rho)}} \frac{1}{g^{\frac{1}{2}} D^{\frac{3}{2}}} \right) \right] = f\left(B \left( \frac{\rho_s - \rho}{\rho} \cdot \frac{D}{R S} \right)\right)$$

or just

$$A \phi_b = f(B \psi) = \phi$$

Einstein settled on  $f(x) = e^{-x}$  for the functional relationship between  $\phi$  and  $\psi$ . He fit  $A$  and  $B$  for natural data, and ended up with  $A=0.465$  and  $B=0.391$ . It turns out that this relationship works pretty well, except at the tail of the graph ( $\phi > 0.4$ ). This may be because  $A$  is no longer a constant, because  $\lambda_0$  isn't constant due to large excess shear. Einstein refined his argument by stating that for high probability the particle wouldn't be expected to settle after only one step ( $\lambda_0 D$ ). He derived a *new*  $\lambda$ :

$$\lambda = \sum_{m=0}^{\infty} (1 - \phi) \phi^{m-1} m \lambda_0 = \frac{\lambda_0}{1 - \phi}$$

which yields curve (2) in the handout.

The whole Einstein model was revised in 1950 by Brown, who fitted a new curve to Einstein's data, based on  $f(x) = x^{-3}$ , rather than  $f(x) = e^{-x}$ . Brown's curve is:

$$\phi = 40 \left( \frac{1}{\psi} \right)^3$$

This curve seems to apply well for low values of  $\psi$  (i.e. high  $\tau_0$  and  $g_b$ ). Einstein's and Brown's curves are often used together, forming the Einstein-Brown bed load equation:

$$\phi = \begin{cases} \frac{1}{0.465} e^{-0.391\psi} & \psi \geq 5.5 \\ 40 \left( \frac{1}{\psi} \right)^3 & \psi \leq 5.5 \end{cases}$$

A few final comments. The Einstein-Brown equation is somewhat more realistic for large  $\psi$  (low  $\tau_0$ ) than DuBoys and other excess shear models, because it shows some transport for  $\tau < \tau_c$ , whereas excess shear models show none. A major drawback of the Einstein-Brown formulation, however, is that there are no bedform effects.